Polarisation effects in spectroscopy of the positronium n = 2 fine structure intervals

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 $^{3}P_{1}(0)$

 $^{3}P_{1}(\pm 1)$



Introduction

- The n = 2 Positronium (Ps, the bound state of an electron and positron) fine structure intervals have been measured several times to test bound-state quantum electrodynamics.[1]
- All previous measurements were performed using waveguides of fixed polarisation.[2,3,4]
- Horn antennas offer a way to change the polarisation of the microwave radiation.

7650h

-35251

-3540

-3555

-3570**-**

-5370

-5385

-5400

--- $J=0, M_{I}=0$

 $- J=2, M_J=\pm 1$

 $J=2, M_{I}=0$

 $\rightarrow \Delta M_I = \pm 1$

 $B_{z}(G)$

100

150

WR-75 Horn Antenna

10 cm

 $\rightarrow \Delta M_1 = 0$

We drove $2^3S_1 \rightarrow 2^{2S+1}P_1$ (S = 0, 1) transitions, known as v_F and v_1 , using a horn antenna to evaluate the effect of polarisation in free-space Ps microwave spectroscopy.

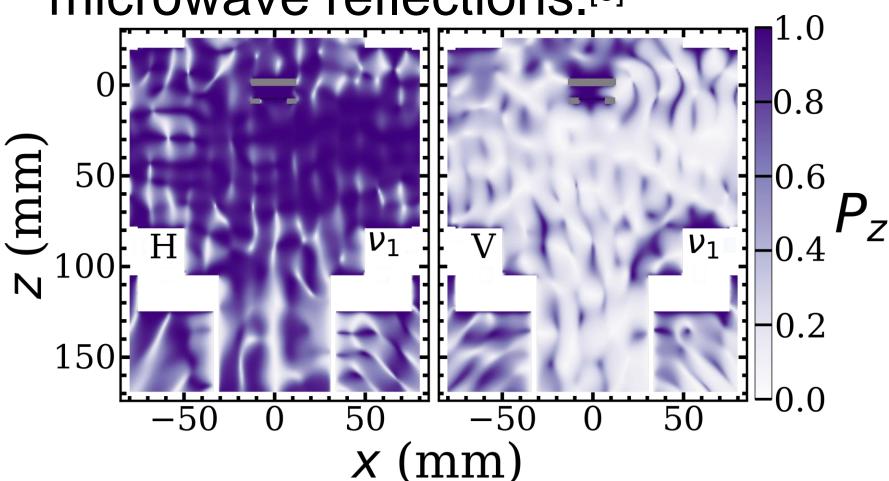
O.02

Zeeman Mixing

- v_F is forbidden in zero field due to charge conjugation symmetry.
- 2¹P₁ levels mix with 2³P₁ levels with the same M₁ in a magnetic field.
- Magnetic field strength (G) The degree of mixing is calculated from the coefficients of the eigenvectors of the full Hamiltonian matrix A_{mi}.

Experimental Methods

- 1³S₁ Ps was made by implanting positron pulses into a mesoporous SiO₂ film in a field of $\langle B_z \rangle = 131$ G.^[6]
- 2³S₁ Ps was made by single photon excitation in an electric field: $1^3S_1 \xrightarrow{\lambda_{UV} = 243 \text{ nm}} 2^3S_1' \xrightarrow{F \approx 2 \to 0 \text{ kV/cm}} 2^3S_1.$ ^[7]
- Retroreflection Time resolved gamma-ray spectroscopy using scintillation detectors allows differentiation of long- and short-lived states.[6]
- A horn antenna emitted radiation parallel to B_{7} when orientated horizontally (H), and orthogonal to B_7 when orientated vertically (V).
- The vacuum chamber was configured to minimise microwave reflections.[5]



Simulations indicate the average unpolarised fraction is 30% for both orientations.

Microwave

Absorbing Foam

The population transfer from the

long-lived 2³S₁ state to the short-lived 2^{2S+1}P₁ state, S_v, was measured as a function of microwave frequency and power.[4]

Fused Silica

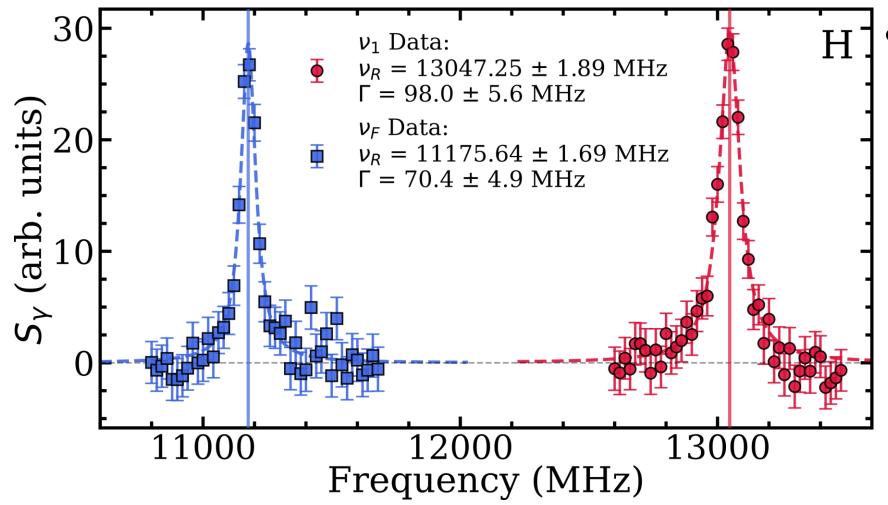
Window

Saturation Measurements

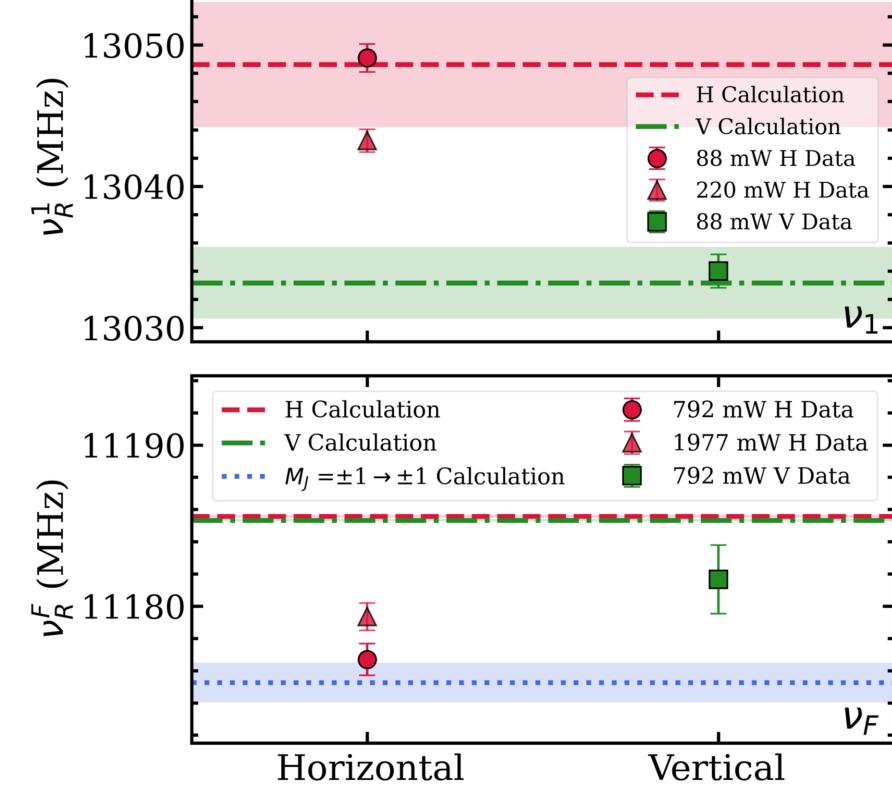
- v₁ saturation data is consistent with all 23S₁ states being depopulated.
- H orientation requires more power to saturate due to 2^3S_1 (0) only being depopulated by low-intensity unpolarised reflections.
- 2000 ν_F does not fully depopulate Horn input power (mW) the 2³S₁ state in V orientation, indicating depopulation of at least one 2³S₁ state is highly limited.

Low line shape power High line shape power Horizontal

Line Shape Measurements



- Line shapes fitted with Lorentzian functions to extract the resonance frequency. Comparison made to average Zeeman shift for each polarisation.
- At low power v₁ data are consistent with transitions driven by fully polarised radiation.
- But v_F is shifted 4.1 σ from theory in the H case. Being consistent with only $2^{3}S_{1}(\pm 1) \rightarrow 2^{1}P_{1}$ (±1) being driven.



High power measurements for H show a shift toward V, indicating low-intensity reflections do become significant.

Conclusions

- We have verified that horn antennas can be used to perform polarisation resolved Ps spectroscopy in freespace. Although near the saturation regime unpolarised reflections can drive additional transitions.
- The v_F saturation and line shape data display unexpected behaviour, likely due to an incomplete model of the Zeeman mixing.
- Ongoing waveguide measurements will further investigate the behaviour of v_F .



[2] S. Hatamian et al., *PRL*, **58**, 1833 (1987)

[3] R. Ley et al., *Hyperfine interactions*, **89**, 327 (1994)

[4] R. E. Sheldon et al., *PRL*, **131**, 043001 (2023)

[5] R. E. Sheldon et al., *PRA*, **107**, 042810 (2023)

[6] B. S. Cooper et al., Rev. Sci. Intrum., 86, 103101 (2015)

[7] A. M. Alonso et al., *PRA*, **95**, 033408 (2017) For more information check out our website: www.antimattergravity.com

